

SWGO perspectives on the GADF

A. Sinha, L. O. Nieto, A. Mitchell, J. Hinton and U. Barres For the SWGO Collaboration

VHE data format call: 29/09/2021

What is SWGO?

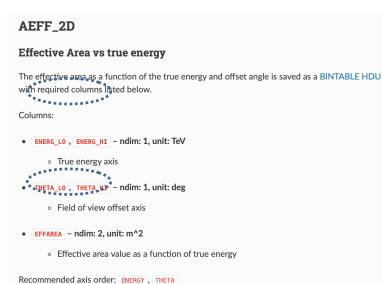
- A proposed Water Cherenkov Detector located in South America
- Energy range ~100s GeV to ~100s TeV
- o High fill-factor core detector + sparse density outer array

IRFs production: The SWGO is very interested to produced GADF compliant IRFs and work with open source tools, but some adaptations to the GADF are requested.

SWGO IRFs - preliminary experiences and remarks

• The GADF format can work for SWGO with some tweaks

- o The GADF definitions are very IACT-y
 - Some inclusive definitions would be nice
 - Definition of an observation
 - ₀ Utility of the obs-index table
 - Mostly time averaged IRFs
 - One CALDB with links to IRFs suffices, hdu-index redundant
 - ₀ The "natural" IRF axes are different for WCD vs IACT
 - ₀ Eg: zenith bins instead of offset bins
 - Possibility to have GADF agnostic to axis names? Eg: bins in (x1, x2, ...)
 - This should be readable by gammapy.maps
- o What is the implication of GTI for WCD?
- o Dealing with different event classes is required
- Will most likely need asymmetric IRF definitions



A DL4 unification

- A typical use case: All sky binned counts and background maps + IRFs
 - Typically called the DL4 level within CTA
 - No DL4 definition exists in the GADF
 - There is an internal gammapy format (Datasets, PSFMap, EDISPMap):
 - Should not be too complicated to generalise it
 - Strong use case within SWGO
 - Should also be useful for joint fitting with other instruments